12. Ozone photochemistry and surface destruction*

The ozone photochemical process in the stratosphere and the ozone destruction by heterogeneous chemical reactions in the planetary boundary layer are parameterized following Cunnold *et al.* (1975), and Schlesinger and Mintz (1979). (see also Schlesinger, 1976)

12.1 Photochemical reactions

The photochemical production and destruction of ozone in the stratosphere is comprised of both the Chapman reactions

$$O_2 + h\nu \xrightarrow{j_1} 2O \tag{12.1}$$

$$O + O_2 + M \xrightarrow{k_1} O_3 + M \tag{12.2}$$

$$O_3 + h\nu \xrightarrow{j_2} O_2 + O \tag{12.3}$$

$$O + O_3 \xrightarrow{k_2} 2O_2 \tag{12.4}$$

and the NO-NO₂ catalytic cycle

$$NO + O_3 \xrightarrow{k_3} NO_2 + O_2 \tag{12.5}$$

$$NO_2 + O \xrightarrow{k_4} NO + O_2$$
 (12.6)

$$NO_2 + h\nu \xrightarrow{j_3} NO + O \tag{12.7}$$

The chemical reaction rates k_i are

$$\begin{aligned} k_1 &= 1.1 \times 10^{-46} exp~(520/T) & m^6 s^{-1}, \\ k_2 &= 1.9 \times 10^{-17} exp~(-2300/T) & m^3 s^{-1}, \\ k_3 &= 2.1 \times 10^{-18} exp~(-1450/T) & m^3 s^{-1}, \\ k_4 &= 9.1 \times 10^{-18} & m^3 s^{-1}, \end{aligned} \tag{12.8}$$

and the photodissociation rate $j_n(p)$ of species n at pressure p is given by

$$j_{n}(p) = \int_{0}^{\infty} \alpha_{n}(\lambda) I_{0}(\lambda) \exp\left[-M_{F} \sum_{m=1}^{2} \alpha_{m}(\lambda) U_{m}(p)\right] d\lambda$$
 (12.9)

Here $\alpha_n(\lambda)$ is the absorption cross section of species n $(1:O_2, 2:O_3, 3:NO_2)$ at wave length

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 λ , $I_0(\lambda)$ is the extraterrestrial monochromatic photon flux per unit wave length. M_F is a magnification factor to account for the departure from the plane parallel atmosphere, which will be defined by (13.51) in Chapter 13, and U_m is the absorber amount of species m (1 : O_2 , 2 : O_3) in a vertical column above pressure p, and the summation extends over all species m. To calculate photodissociation rates $j_n(p)$, the values of $\alpha_1(\lambda)$, $\alpha_2(\lambda)$ and $I_0(\lambda)$ are taken from Ackermann (1971), Kockarts (1971), and $\alpha_3(\lambda)$ from the data of The Natural Stratoshere of 1974, CIAP Monograph 1. Currently $j_n(p)$ is computed by a linear bivariate interpolation of precomputed values of $j_n(p)$ as a function of U_1 and U_2 to save CPU time.

12.2 Governing equations for the photochemical reactions

The equations governing the time change rates of the concentration of O, O_2 , O_3 , NO, and NO_2 given by reactions (12.1)—(12.7) are

$$\frac{\partial(O)}{\partial t} = 2j_1(O_2) - k_1(O)(O_2)(M) + j_2(O_3) - k_2(O)(O_3) - k_4(O)(NO_2) + j_3(NO_2),$$
(12.10)

$$\frac{\partial(O_2)}{\partial t} = -j_1(O_2) - k_1(O)(O_2)(M) + j_2(O_3) + 2k_2(O)(O_3) + k_3(NO)(O_3) + k_4(NO_2)(O),$$
(12.11)

$$\frac{\partial(O_3)}{\partial t} = k_1(O)(O_2)(M) - j_2(O_3) - k_2(O)(O_3) - k_3(NO)(O_3), \tag{12.12}$$

$$\frac{\partial (NO)}{\partial t} = -k_3(NO)(O_3) + k_4(NO_2)(O) + j_3(NO_2), \tag{12.13}$$

$$\frac{\partial (NO_2)}{\partial t} = k_3(NO)(O_3) - k_4(NO_2)(O) - j_3(NO_2), \tag{12.14}$$

where (X) denotes the concentration of species X in molecules m⁻³ and M represents that of air.

The equilibrium concentrations of atomic oxygen O and nitric monooxide NO are calculated by (12.10) and (12.13),

$$(O)_{e} = \frac{2j_{1}(O_{2}) + j_{2}(O_{3}) + j_{3}(NO_{2})}{k_{1}(O_{2})(M) + k_{2}(O_{3}) + k_{4}(NO_{2})},$$
(12.15)

$$(NO)_{e} = \frac{k_{4}(NO_{2})(O) + j_{3}(NO_{2})}{k_{3}(O_{3})},$$
(12.16)

the relaxation times are approximately given by $\tau_0 \approx (k_1(O_2)(M))^{-1}$ and $\tau_{No} \approx (k_3(O_3))^{-1}$; their representative values are order 1 minute and 10 minutes, respectively. Then, it can be assumed that O and NO are in equilibrium with other constituents, and we can approximate $(O)_e$ and $(NO)_e$ instead of (O) and (NO).

Substituting (12.15) and (12.16) into (12.12) gives

$$\frac{\partial(O_3)}{\partial t} = A - B(O_3) - C(O_3)^2 \tag{12.17}$$

where

$$A = 2\frac{k_1j_1(O_2)^2(M) - k_4j_1(O_2)(NO_2) - k_4j_3(NO_2)^2}{k_1(O_2)(M) + k_2(O_3) + k_4(NO_2)}$$

$$B = 2 \frac{k_2 j_1(O_2) + (k_2 j_3 + k_4 j_2)(NO_2)}{k_1(O_2)(M) + k_2(O_3) + k_4(NO_2)}$$

$$C = 2 \frac{k_2 j_2}{k_1(O_2)(M) + k_2(O_3) + k_4(NO_2)}$$

 (O_2) can be given as $(O_2) = 0.21(M)$, and (NO_2) is prescribed to vary only in the vertical with the relation

$$(NO2)(z) = \beta(NO2)MCE Iroy(z)$$
(12.18)

where $(NO_2)_{MCEIroy}(z)$ is the one-dimensional profile calculated by McElroy *et al.* (1974), and β is an empirical constant to adjust the simulated O_3 mixing ratio to the observed value at 10 mb in the tropics. Currently β is set to 1.62.

12.3 Vertical distribution of absorber amounts

The integrated absorber amount of species m in a vertical column above pressure p is defined by

$$U_{m}(p) = \frac{1}{g} \int_{0}^{p} (-)_{m}(p) \frac{dp}{\rho}$$
 (12.19)

where ()_m is the concentration of species m in molecules m^{-3} . We assume that molecular oxygen is well-mixed throughout the atmosphere, $(O_2) = 0.21(M)$; then for molecular oxygen

$$U_1(p) = \frac{0.21 \,\alpha p}{g \,m} \tag{12.20}$$

where α is Avogadro's number, and m is the molecular weight of air. For ozone, it is assumed that the number density $(O_3(z))$ above the midlevel of layer 1 decays exponentially

with altitude following Krueger (1973). Thus

$$(O_3(z)) = (O_3(z_1)) \exp(-\frac{z - z_1}{H}) \quad z \ge z_1$$
 (12.21)

where $(O_3(z_1))$ is the GCM predicted ozone number density for layer 1, and H is the scale height for ozone (=4.45km). Then (12.19) can be written as

$$U_{2}(p) = \int_{z_{1}}^{\infty} (O_{3}(z)) dz + \frac{\alpha}{g \mathcal{M}_{03}} \int_{P_{1}}^{P} q_{03} (p) dp$$
 (12.22)

where \mathcal{M}_{03} is the molecular weight of ozone.

12.4 Ozone destruction at the earth's surface

Ozone is destroyed at the earth's surface by heterogeneous chemical reactions. The rate of destruction D_{03} may be expressed as

$$D_{03} = \rho K (q_{03})_s \tag{12.23}$$

where ρ is the air density, K the reaction rate constant, $(q_{03})_s$ the surface ozone mixing ratio. We assume that the destruction D_{03} is approximately compensated by the downward ozone flux at the top of the planetary boundary layer (PBL), $(F_{03})_B$ which may be approximated as

$$(F_{03})_{B} = \rho D \frac{\partial q_{03}}{\partial Z} = \rho D \frac{(q_{03})_{LM} - (q_{03})_{S}}{Z_{LM} - Z_{S}}$$
(12.24)

where D is the eddy diffusivity at the top of the PBL, Z the altitude, and subscript LM denotes the midlevel of the lowest layer. $(q_{03})_s$ is determined by equating (12.23) with (12.24) as

$$(q_{03})_{s} = \frac{D}{D + K(Z_{1M} - Z_{s})}(q_{03})_{LM}$$
(12.25)

The constants K and D are currently assigned to 0.0008 m sec⁻¹ and 10 m²sec⁻¹ after Cunnold *et al.* (1975).